

SIZE:

30-meters
in diameter

MIRROR:

492
segments

of reflective glass pieced
together to form one
giant primary mirror

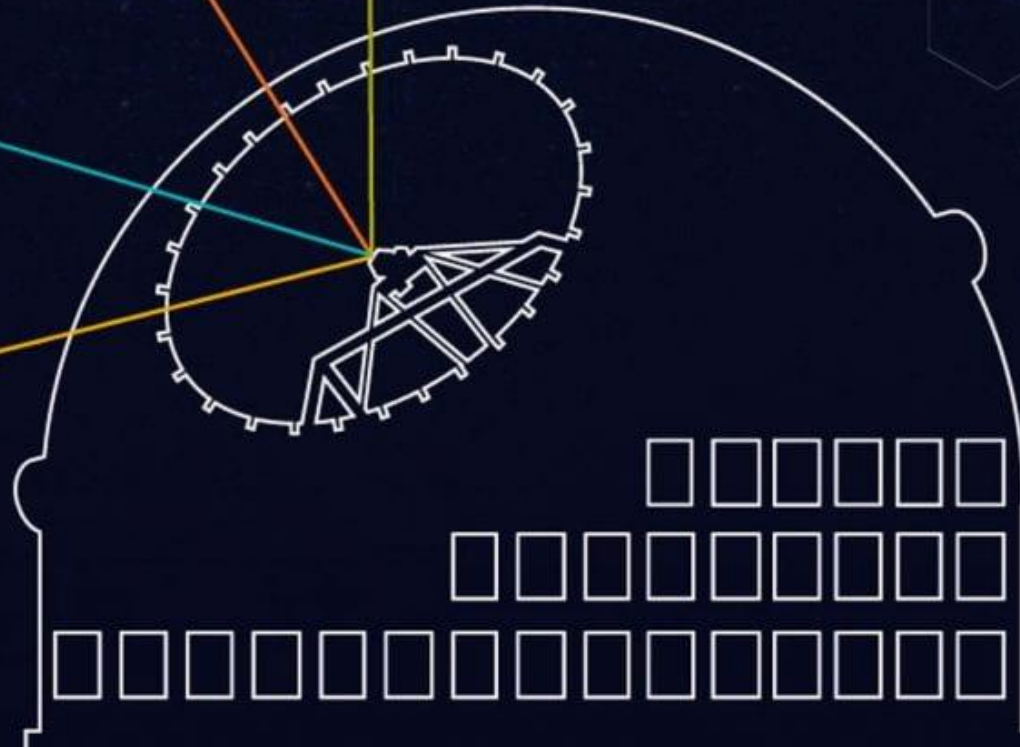
TOTAL COLLECTING AREA:

655
square meters

RESOLUTION:

12 times
sharper

than that of the
Hubble Space Telescope



TMTで紐解く
惑星系の起源

長谷川靖紘

ジェット推進研究所
カリフォルニア工科大学



Jet Propulsion Laboratory
California Institute of Technology

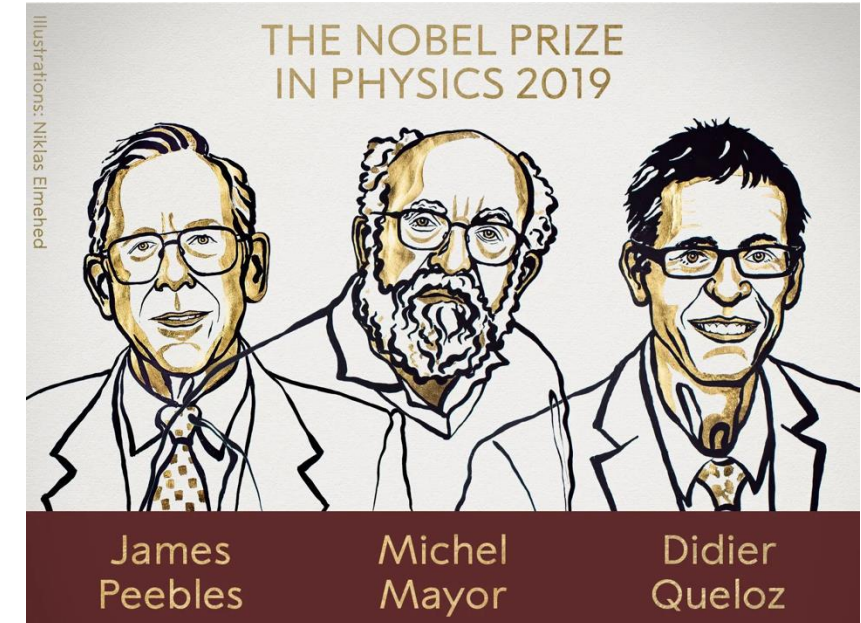
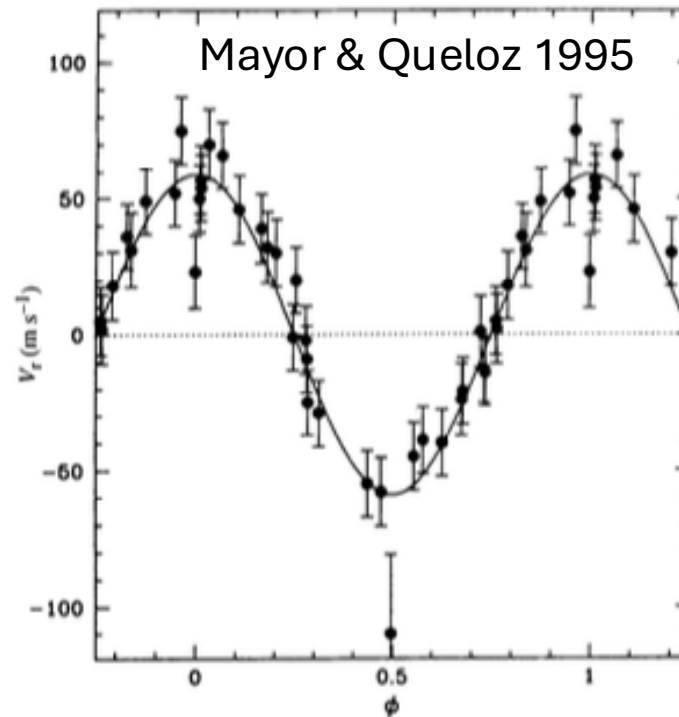
Why Do We Care?

How many planets are out there?
Is the solar system unique?
Are we alone?

Discoveries of Extrasolar Planets (Exoplanets)



©JPL/NASA



Adapted from <https://www.bloomberg.com/news/articles/2019-10-08/peebles-mayor-and-queloz-share-2019-nobel-prize-in-physics>

The first discovery of a gas giant planet orbiting around other Sun-like star

The radial velocity (i.e., wobble) of a star caused by a planet was observed

The unexpected presence of a hot Jupiter surprised the community!

Fundamental questions:

When, Where, and How Do Planets Form?

"Classical" Picture of Planet Formation

e.g., Hayashi 1981

Natural outcome
of star formation

Minimum-mass
Solar nebula

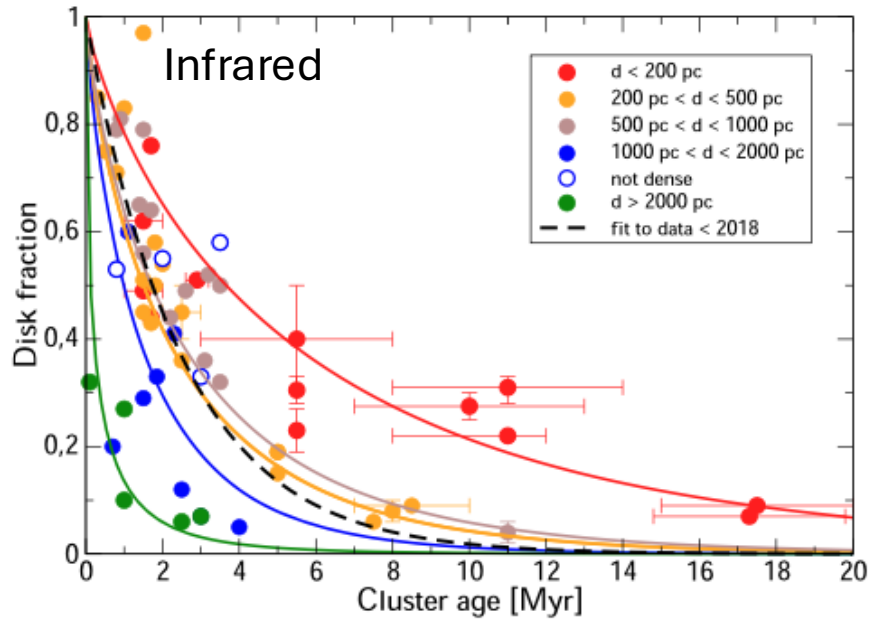
In-situ
formation

Dynamically
inactive

NEBULAR HYPOTHESIS

State-of-the Art: Disk Observations

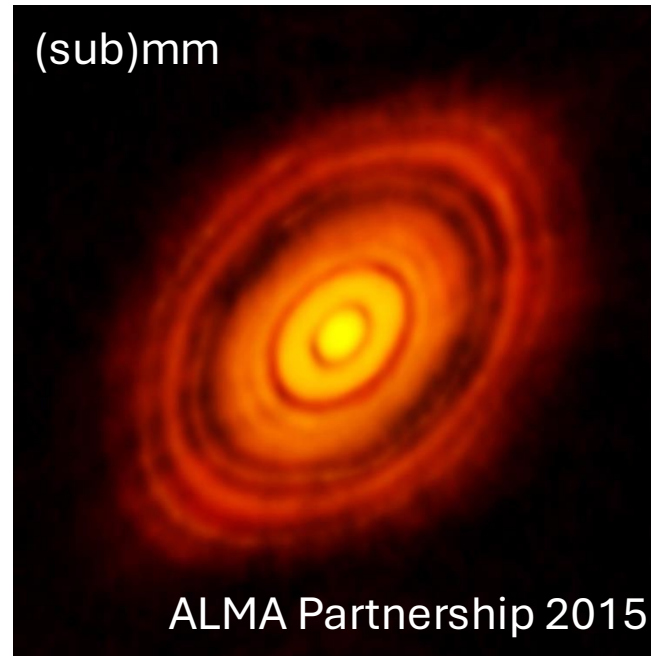
When



e.g., Pfalzner et al 2022

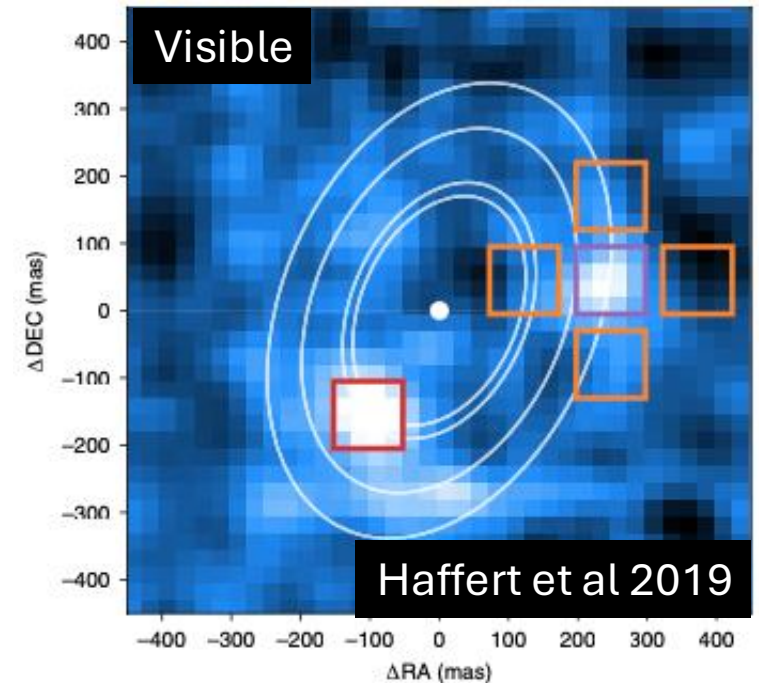
The lifetime of protoplanetary disks is typically about 1-10 Myr

Where



Gaps can be opened by embedded, forming planets

How

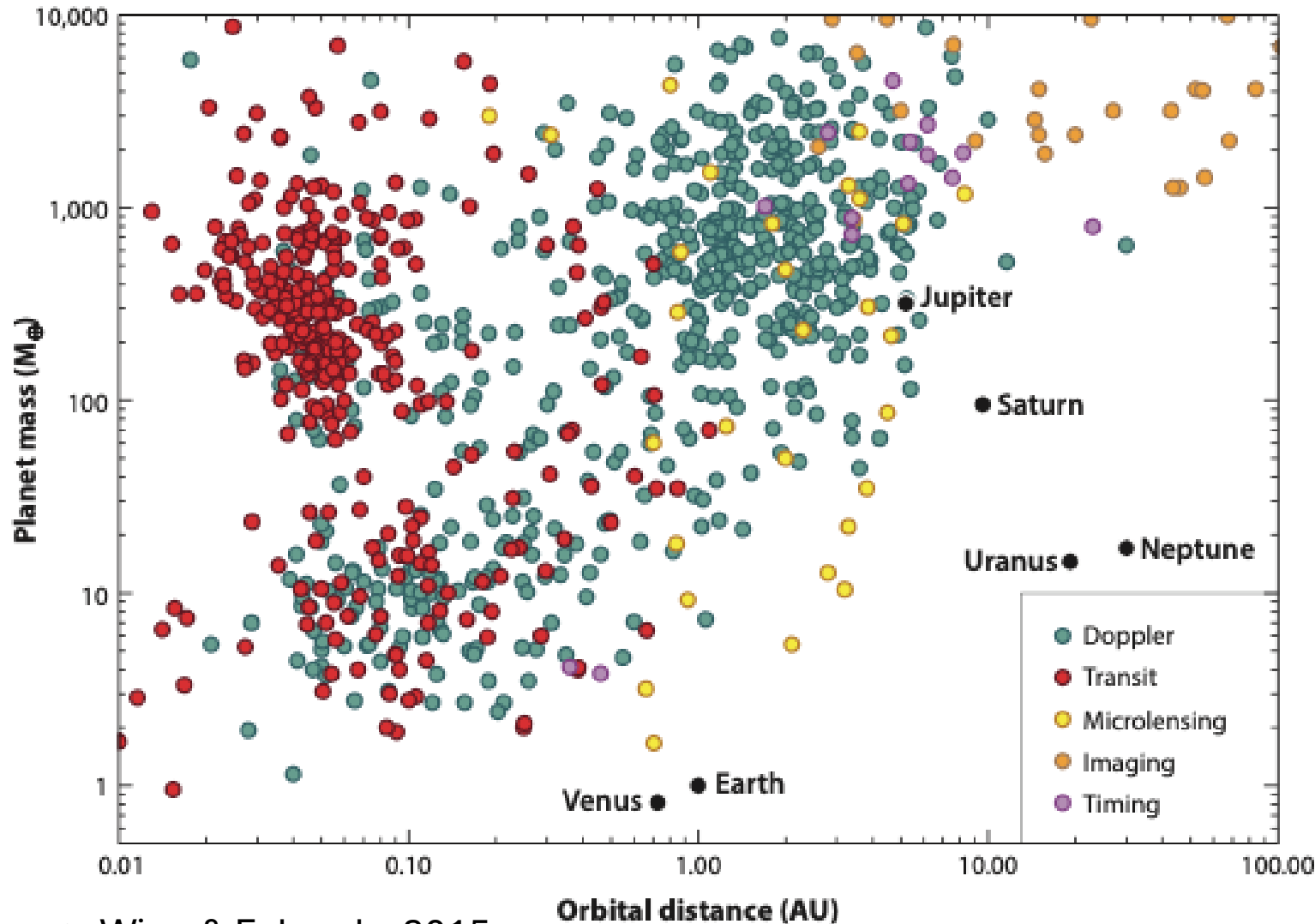


Halpha emission comes from accreting giant planets

State-of-the Art: Exoplanet Population

Where

How



e.g., Winn & Fabrycky 2015

Exoplanets are everywhere

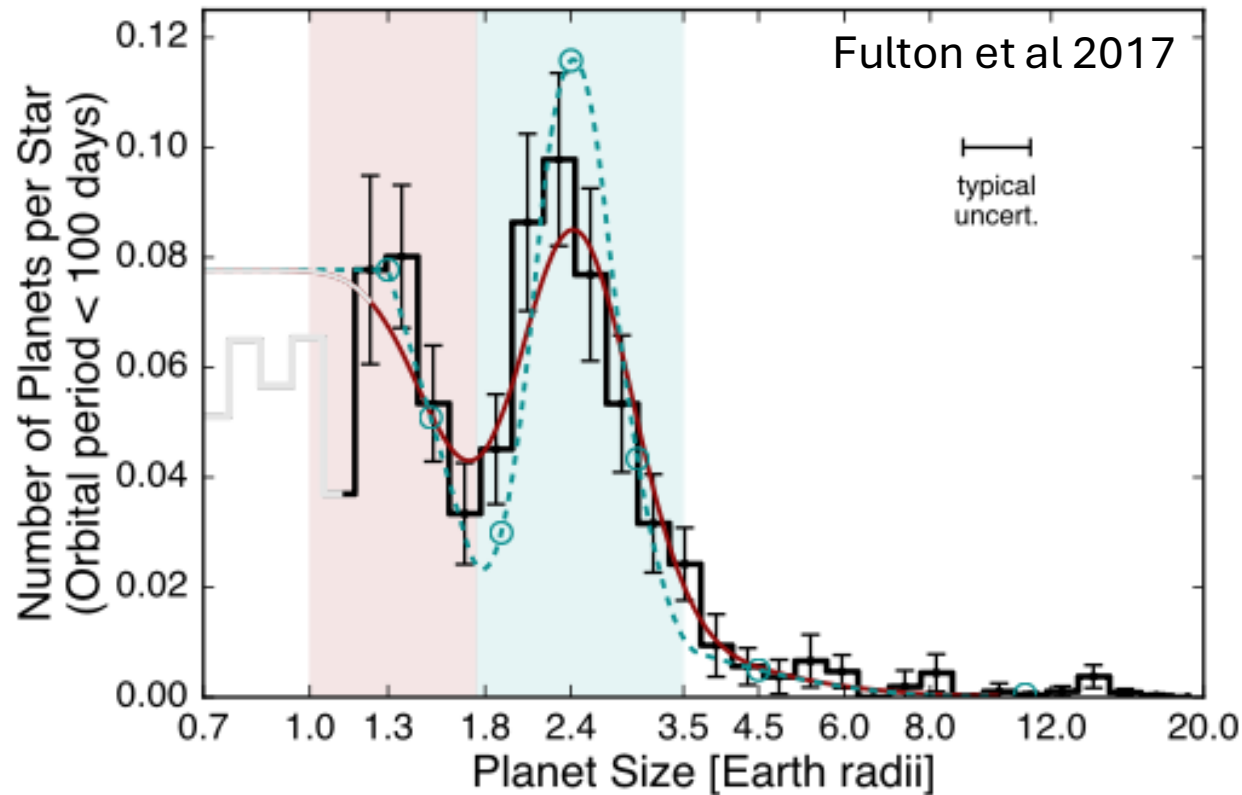
Planet formation is ubiquitous

Accretion of gas and solids
leads to a wide mass range

Orbital evolution
plays an important role

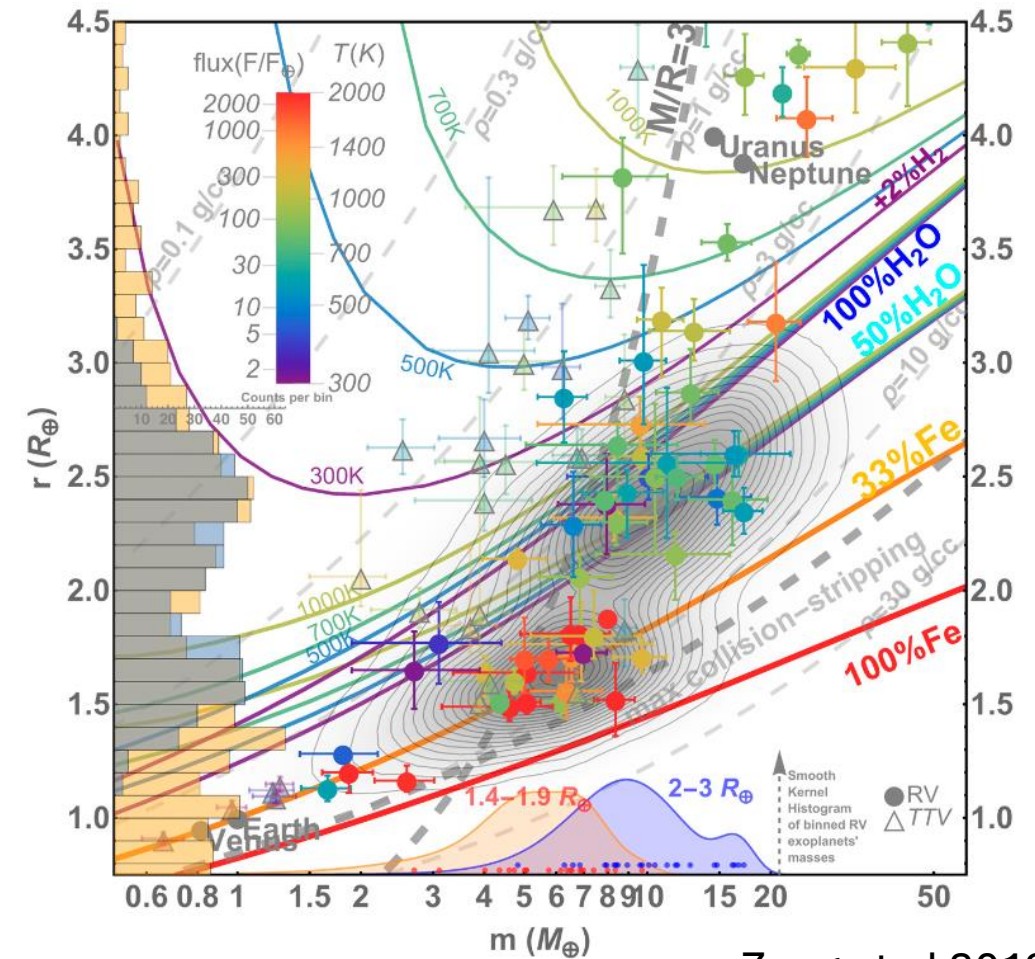
State-of-the Art: Exoplanet Composition

How



Planets with radius of $< 1.8 R_{\text{Earth}}$
are likely rocky

Where & How



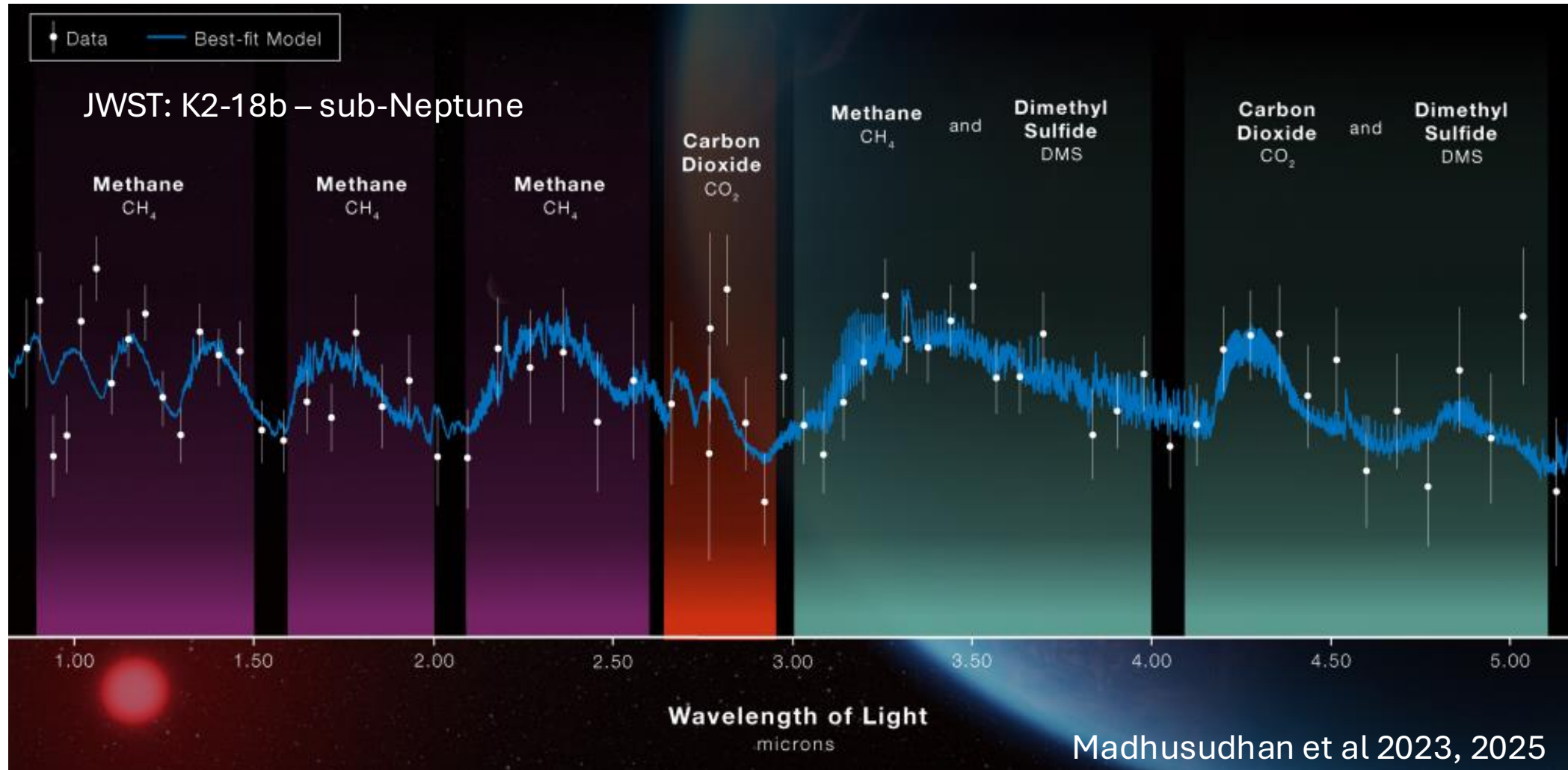
Zeng et al 2019

Some sub-Neptune planets are volatile rich

State-of-the Art: Exoplanet Climate & Habitability

How

Where



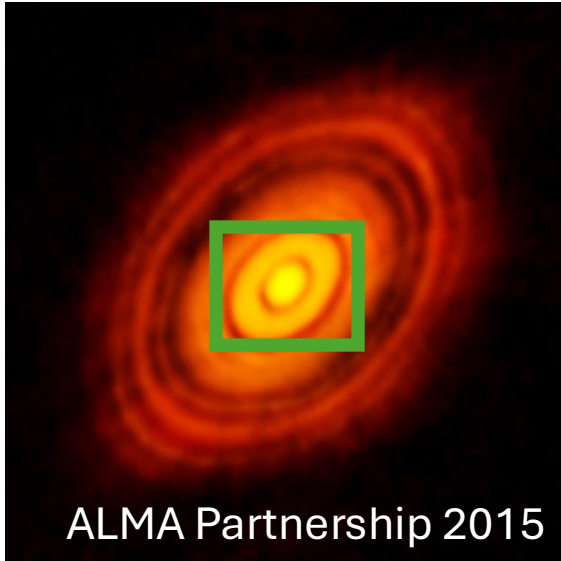
Characterization of exoplanet atmospheres can infer climate and even habitability of exoplanets

Fundamental questions:

When, Where, and How Do Planets Form?

Beyond State-of-the Art: Discovery-Based

Disks

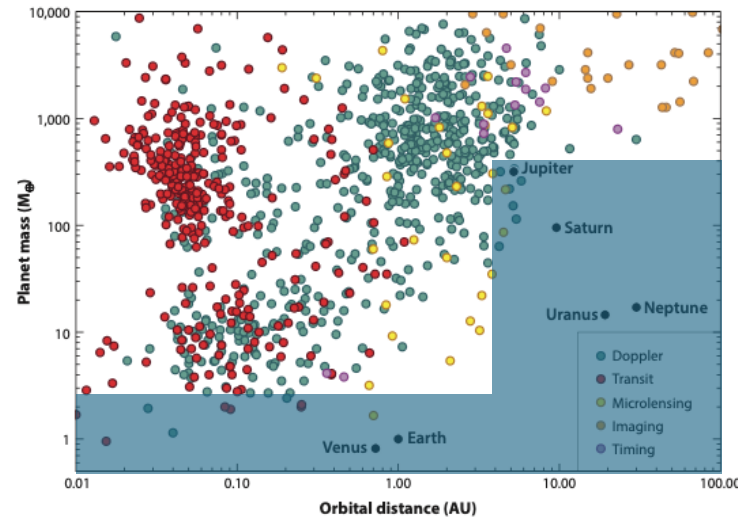


Resolve inner disks

Gas kinematics

Multi-wavelength

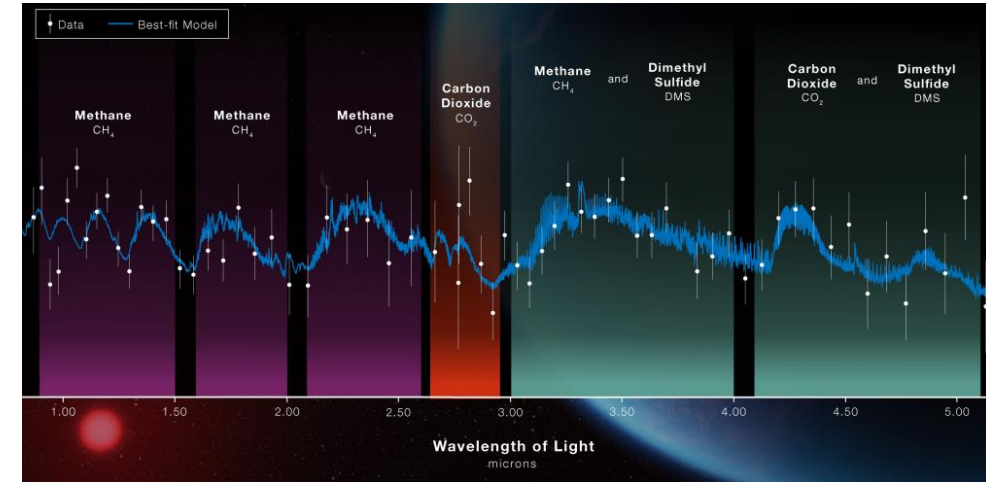
Population



e.g., Winn & Fabrycky 2015

Low-mass planets including habitable planets

Characterization



Madhusudhan et al 2023, 2025

Distant planets

Young planets

Multi-wavelength

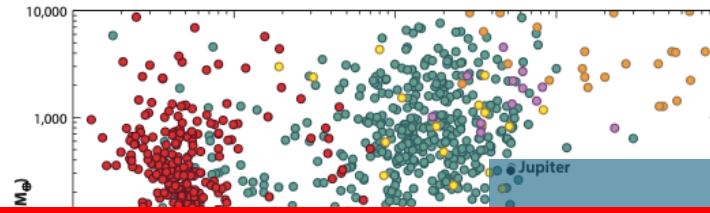
More atoms/molecules

Beyond State-of-the Art: Discovery-Based

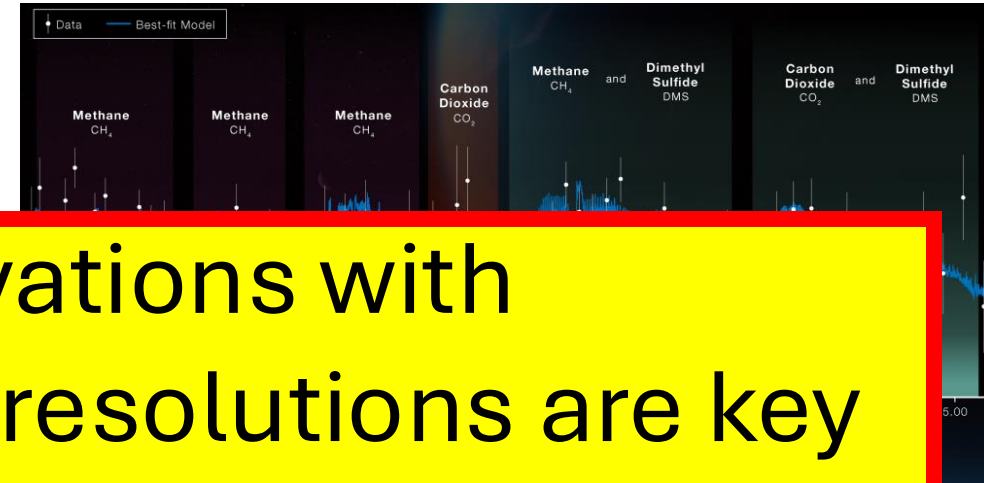
Disks



Population



Characterization



High sensitivity observations with high spatial and high spectral resolutions are key while setting more specific goals is preferred

Resolve inner disks

Low-mass planets including habitable planets

Gas kinematics

Distant planets

Multi-wavelength

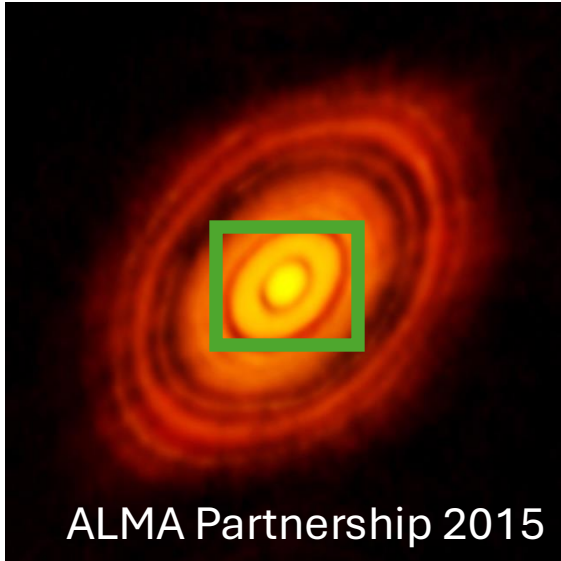
Multi-wavelength

Young planets

More atoms/molecules

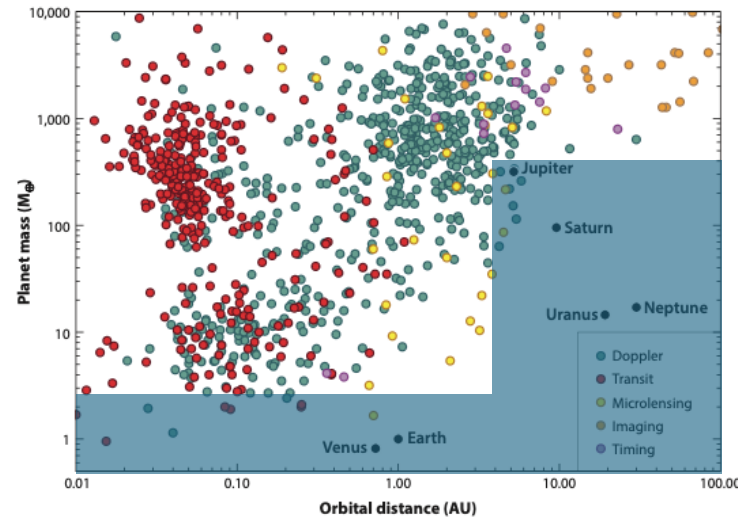
Beyond State-of-the Art: Hypothesis-Based

Disks



Are inner disk properties set by disk winds or turbulence?

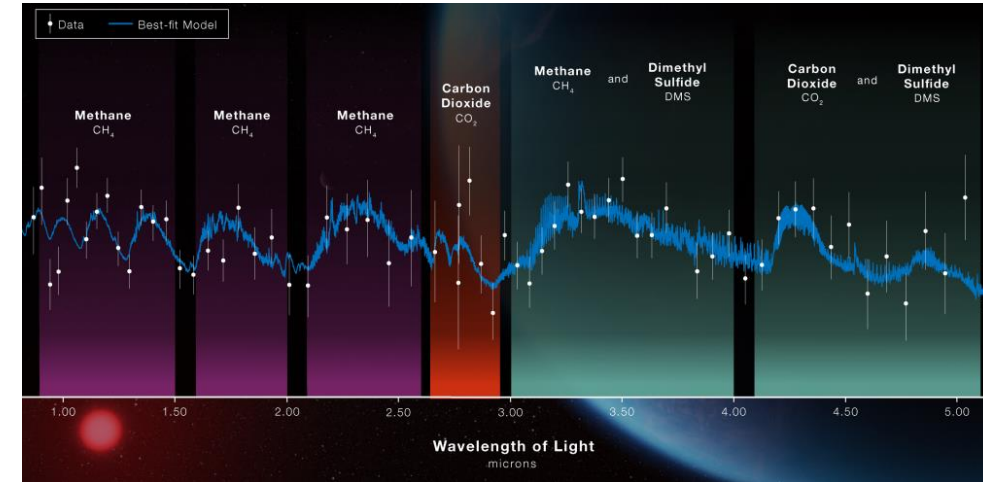
Population



e.g., Winn & Fabrycky 2015

Did Low-mass planets including habitable planets form by in-situ or gas-induced migration?

Characterization

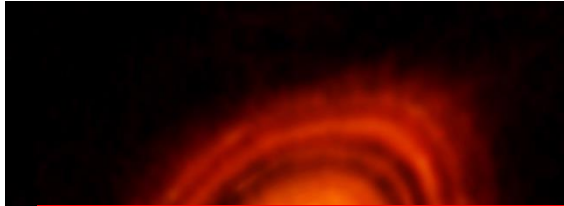


Madhusudhan et al 2023, 2025

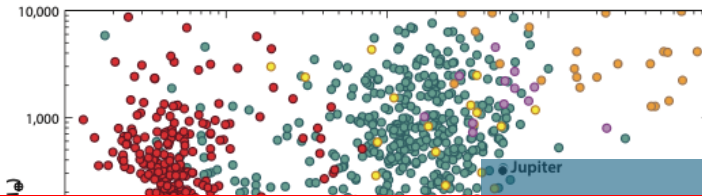
Is the current properties of molecular abundance determined by equilibrium or non-equilibrium chemistry?

Beyond State-of-the Art: Hypothesis-Based

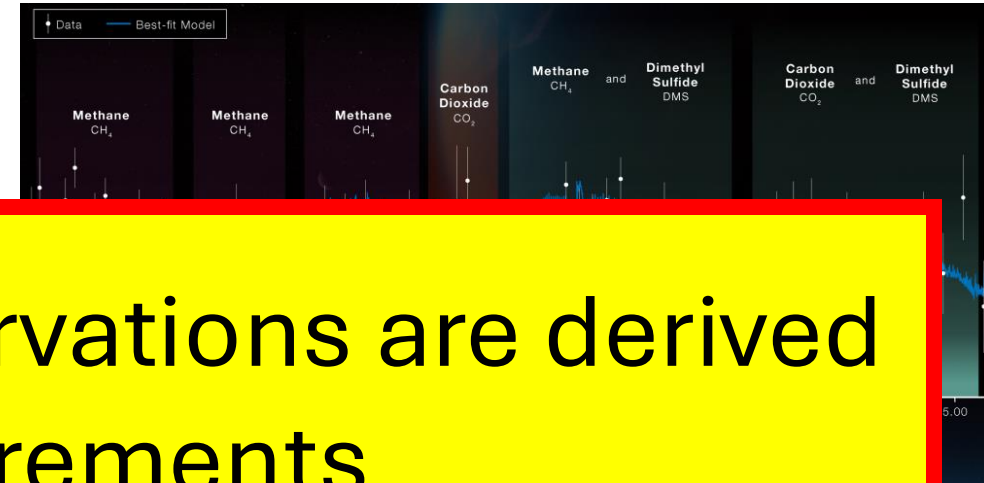
Disks



Population



Characterization



The specs of telescopes/observations are derived from science requirements

Are inner disk properties set by disk winds or turbulence?

Did Low-mass planets including habitable planets form by in-situ or gas-induced migration?

Is the current properties of molecular abundance determined by equilibrium or non-equilibrium chemistry?

Key Science Cases for TMT

Transit!

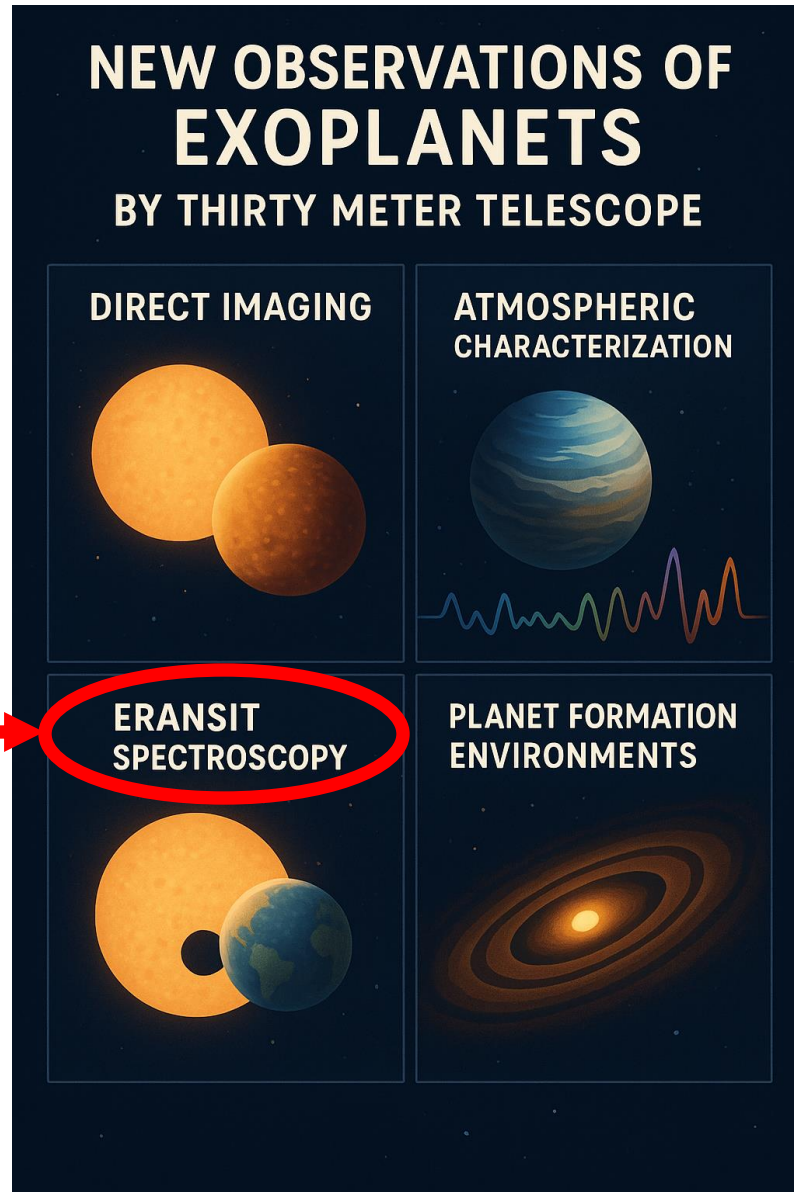
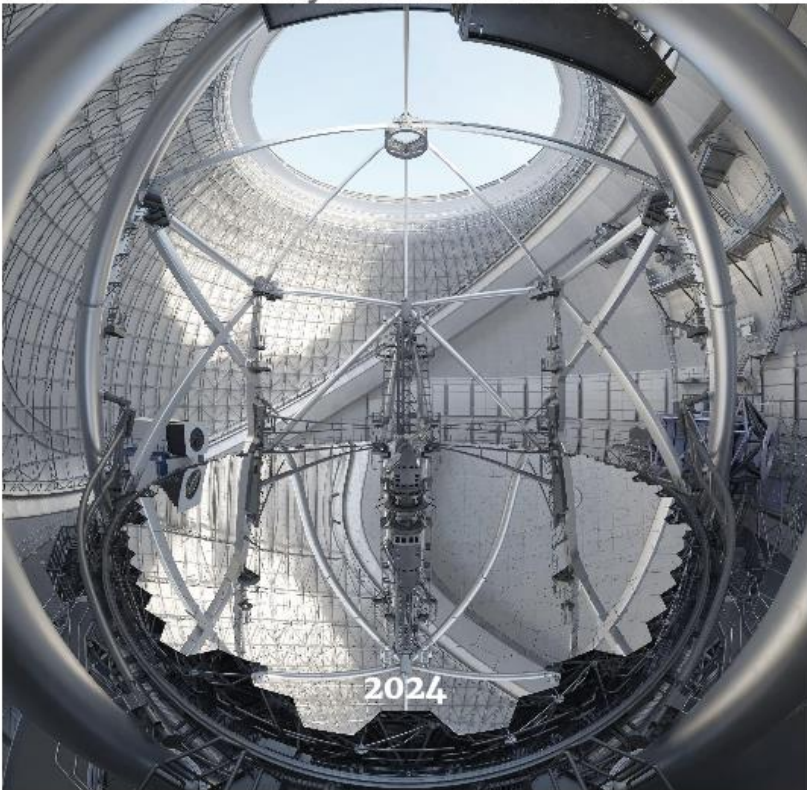


Image created by ChatGPT!

Science Requirements for TMT

Science cases from the community

Thirty Meter Telescope International
Observatory Detailed Science Case



TIO.PSC.TEC.07.007.CCR04

<https://www.tmt.org/download/Document/10/original>

Planet formation &
exoplanet science cases

Imaging inner (10 au) disks

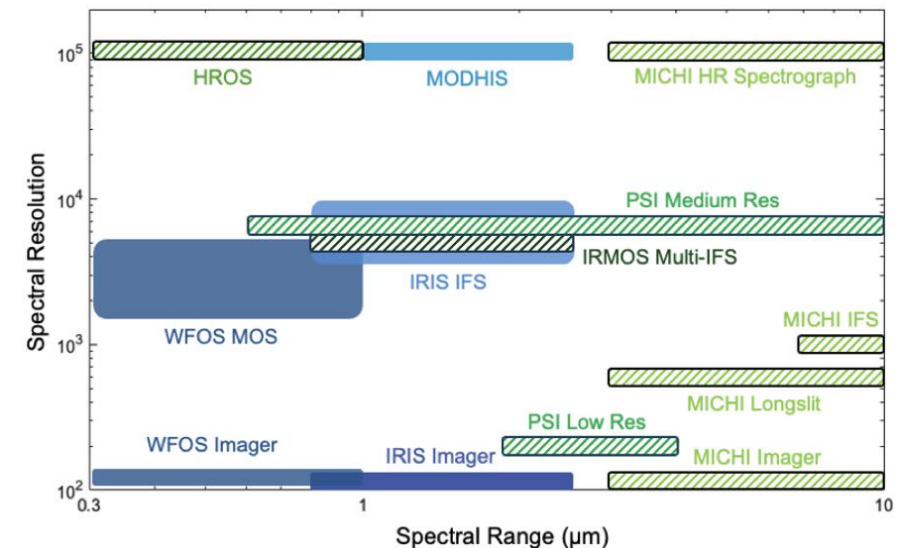
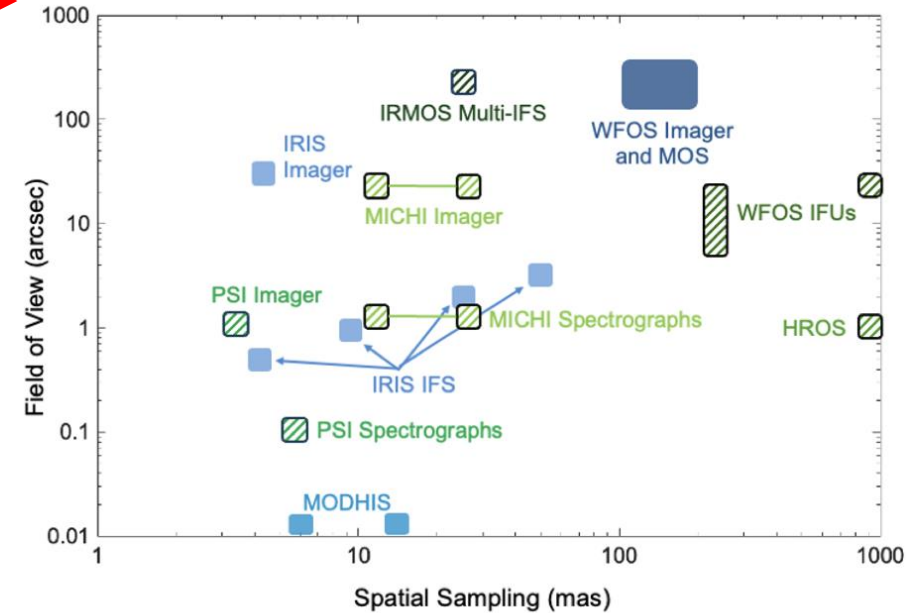
Spectroscopy of disk gas

Imaging & spectroscopy of
young giant planets

RV detections and transit
spectroscopy of (super)Earths
in habitable zones

RV detections of exomoons

Instrument specs



Some numbers...

First Light Instruments

MODHIS

10 mas @ 150 pc \Rightarrow 1.5 au

$R \sim 10^5 \Rightarrow \Delta v \sim 3 \text{ km/s}$

1 micron \Rightarrow 3000K

IRIS

1 as FoV @ 150 pc \Rightarrow 150 au

$R \sim 10^4 \Rightarrow \Delta v \sim 30 \text{ km/s}$

1 micron \Rightarrow 3000K

First Decade Instruments

PSI

5 mas @ 10 pc \Rightarrow 0.05 au

$R < 10^5 \Rightarrow \Delta v \sim 3 \text{ km/s}$

Contrast ratio $\sim 10^{-8}$ - 10^{-9}

MICHI

10 mas @ 150 pc \Rightarrow 1.5 au

$R < 10^5 \Rightarrow \Delta v \sim 3 \text{ km/s}$

10 micron \Rightarrow 300 K

Key physical quantities

$v_{\text{Kep}} @ 1 \text{ au} \Rightarrow 30 \text{ km/s}$ RV of 10 cm/s can be achieved by $R \sim 10^5$ Contrast ratio of Jupiter $\sim 10^{-9}$

Synergy with Other Telescopes

Synergy	Ch#3	Ch#4	Ch#5	Ch#6	Ch#7	Ch#8	Ch#9	Ch#10	Ch#11
	Fund. Physics and Cosmology	Early Universe	Galaxy Formation and the IGM	SMBH	MW and Nearby Galaxies	Birth & Early Lives of Stars & Planets	Time-Domain Science	Exoplanets	Our Solar System
ALMA		☑	☑ ☑	☑ ☑		☑ ☑	☑ ☑	☑ ☑	
GAIA	☑ ☑				☑ ☑	☑ ☑	☑ ☑	☑ ☑	
JWST	☑ ☑	☑ ☑	☑ ☑	☑ ☑	☑ ☑	☑ ☑	☑ ☑	☑ ☑	☑ ☑
Roman Space Telescope	☑ ☑	☑ ☑	☑	☑	☑	☑	☑ ☑	☑ ☑	
HWO	☑	☑	☑	☑	☑	☑	☑	☑ ☑	☑

<https://www.tmt.org/download/Document/10/original>

TMT will be very versatile and provide complementary observations

Proposing science cases unique to TMT and a next-generation of instruments can enhance the importance of TMT

TMT-ACCESS Workshop

What is the TMT-ACCESS?

新たな TMT サイエンスケースの創造 と TMT 次世代装置の提案
を目的とした 若手中心の TMT Workshop Series

分野横断型

ディスカッション重視

若手中心

サイエンティスト
— エンジニアの
相互理解促進



TMT-ACCESS Workshop

Former workshops

Our Activities: TMT-ACCESS

4



First TMT-ACCESS

日時 2023年 9月
場所 TIO Office (アメリカ/パサデナ)
テーマ TMT が拓く次世代のサイエンス
講義 TMT/ELT のサイエンス、TMT の装置、
Hawaiian engagement、研究分野毎のレビュー
見学 TMT ラボ、JPL



梅畑 豪紀 (名古屋大学)(CSU Northridge) 鵜山 太智 (信州大学)
砂山 朋美 (ASIAA) 高橋 葵 (JAXA)
中本 崇志 (NAOJ) 長谷川 靖紘 (JPL) 百瀬 莉恵子 (Carnegie)



TMT-ACCESS 2024

日時 2024年 6月
場所 東北大学 (仙台)
テーマ 日本がリードする20年後の光赤外線天文学
講義 装置開発、研究分野毎のレビュー
見学 天文学専攻・地球物理学専攻ラボ



小野里 宏樹 (NAOJ) 木村 成生 (東北大学) 久保 真理子 (東北大学) 鈴木 竜二 (NAOJ)
田畑 陽久 (ISAS/JAXA) 富田 賢吾 (東北大学)

TMT-ACCESS Workshop

Next 2025

TMT-ACCESS

TMT eArly Career Centered,
Engineers-Scientists Synergy

will be at

NAOJ, July 16–18

tmt-access@ml.nao.ac.jp



鵜山 太智
(CSU Northridge)



小野里 宏樹
(NAOJ)



衣川 智弥
(信州大学)



鈴木 竜二
(NAOJ)



瀧本 幸司
(ISAS/JAXA)



田畑 陽久
(ISAS/JAXA)



野津 翔太
(東京大学)



長谷川 靖紘
(JPL)



百瀬 莉恵子
(Carnegie)



森 万由子
(ABC/NAOJ)



米田 謙太
(NAOJ)

Summary

- Planet formation and exoplanets are timely in astrophysics today
- Tremendous amounts of progresses have been made
- The origins of planets – especially habitable planets – remain unclear
- High sensitive observations with high spatial and high spectral resolutions can advance the fields significantly
- TMT will be very versatile and can play such a role
- Due to complementarity of TMT observations, proposing science cases unique to TMT and a next-generation of instruments would be key
- Participation of the TMT-ACCESS workshop is a great opportunity and highly encouraged